

E.M. Ovsyuk¹, O.V. Veko**SPIN 1/2 PARTICLE IN THE FIELD OF THE DIRAC STRING ON THE
BACKGROUND OF ANTI DE SITTER SPACE-TIME**

Mozyr State Pedagogical University named after I.P. Shamyakin

Abstract

The Dirac monopole string is specified for anti de Sitter cosmological model. Dirac equation for spin 1/2 particle in presence of this monopole has been examined on the background of anti de Sitter space-time in static coordinates. Instead of spinor monopole harmonics, the technique of Wigner D -functions is used. After separation of the variables radial equations have been solved exactly in terms of hypergeometric functions. The complete set of spinor wave solutions $\Psi_{\epsilon,j,m,\lambda}(t, r, \theta, \phi)$ has been constructed, the most attention is given to treating the states of minimal values for total moment quantum number j_{min} . At all values of j , the energy spectrum is discrete.

PACS numbers: 11.10.Cd, 04.20.Gz

1 Introduction

De Sitter and anti de Sitter geometrical models are given steady attention in the context of developing quantum theory in a curved space-time – for instance, see in [1]. In particular, the problem of description of the particles with different spins on these curved backgrounds has a long history – see [2–34]. Here we will be interested mostly in treating the Dirac equation in de Sitter model. In the present paper, the influence of the Dirac monopole string on the spin 1/2 particle in the anti de Sitter cosmological model is investigated. Instead of spinor monopole harmonics, the technique of Wigner D -functions is used. After separation of the variables radial equation have been solved exactly in terms of hypergeometric functions. The complete set of spinor wave solutions $\Psi_{\epsilon,j,m,\lambda}(t, r, \theta, \phi)$ has been constructed. Special attention is given to treating the states of minimal values for total moment quantum number j_{min} , these states turn to be much more complicated than in the flat Minkowski space. At all values of j , the energy spectrum is discrete.

2 Dirac particle in the anti de Sitter space

The Dirac equation (the notation according to [39] is used)

$$\left[i\gamma^c (e_{(c)}^\alpha \partial_\alpha + \frac{1}{2}\sigma^{ab}\gamma_{abc}) - M \right] \Psi = 0 \quad (1)$$

¹e.ovsyuk@mail.ru

in static coordinates and tetrad of the anti de Sitter space-time

$$\begin{aligned}
dS^2 &= \Phi dt^2 - \frac{dr^2}{\Phi} - r^2(d\theta^2 + \sin^2 \theta d\phi^2), \quad \Phi = 1 + r^2, \\
e_{(0)}^\alpha &= \left(\frac{1}{\sqrt{\Phi}}, 0, 0, 0\right), \quad e_{(3)}^\alpha = (0, \sqrt{\Phi}, 0, 0), \\
e_{(1)}^\alpha &= \left(0, 0, \frac{1}{r}, 0\right), \quad e_{(2)}^\alpha = \left(1, 0, 0, \frac{1}{r \sin \theta}\right), \\
\gamma_{030} &= \frac{\Phi'}{2\sqrt{\Phi}}, \quad \gamma_{311} = \frac{\sqrt{\Phi}}{r}, \quad \gamma_{322} = \frac{\sqrt{\Phi}}{r}, \quad \gamma_{122} = \frac{\cos \theta}{r \sin \theta},
\end{aligned} \tag{2}$$

takes the form

$$\begin{aligned}
&\left[i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \left(\gamma^3 \partial_r + \frac{\gamma^1 \sigma^{31} + \gamma^2 \sigma^{32}}{r} + \right. \right. \\
&\quad \left. \left. \frac{\Phi'}{2\Phi} \gamma^0 \sigma^{03} \right) + \frac{1}{r} \Sigma_{\theta, \phi} - M \right] \Psi(x) = 0,
\end{aligned} \tag{3}$$

where

$$\Sigma_{\theta, \phi} = i \gamma^1 \partial_\theta + \gamma^2 \frac{i \partial + i \sigma^{12} \cos \theta}{\sin \theta}.$$

Eq. (3) reads

$$\left[i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \gamma^3 \left(\partial_r + \frac{1}{r} + \frac{\Phi'}{4\Phi} \right) + \frac{1}{r} \Sigma_{\theta, \phi} - M \right] \Psi(x) = 0. \tag{4}$$

From (4), with the substitution $\Psi(x) = r^{-1} \Phi^{-1/4} \psi(x)$, we get

$$\left(i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \gamma^3 \partial_r + \frac{1}{r} \Sigma_{\theta, \phi} - M \right) \psi(x) = 0. \tag{5}$$

Below the spinor basis will be used

$$\gamma^0 = \begin{vmatrix} 0 & I \\ I & 0 \end{vmatrix}, \quad \gamma^j = \begin{vmatrix} 0 & -\sigma_j \\ \sigma_j & 0 \end{vmatrix}, \quad i\sigma^{12} = \begin{vmatrix} \sigma_3 & 0 \\ 0 & \sigma_3 \end{vmatrix}.$$

3 Separation of the variables

Let us start with the monopole Abelian potential in the Schwinger's form [35] in flat Minkowski space

$$A^a(x) = (A^0, A^i) = \left(0, g \frac{(\vec{r} \times \vec{n}) \cdot (\vec{r} \vec{n})}{r(r^2 - (\vec{r} \vec{n})^2)} \right). \tag{6}$$

Specifying $\vec{n} = (0, 0, 1)$ and translating the A_α to the spherical coordinates, we get

$$A_0 = 0, \quad A_r = 0, \quad A_\theta = 0, \quad A_\phi = g \cos \theta. \tag{7}$$

It is easily verified that this potential A_ϕ obeys Maxwell equations in anti de Sitter space

$$\begin{aligned} \frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^\alpha} \sqrt{-g} F^{\alpha\beta} &= 0, \sqrt{-g} = r^2 \sin \theta, \\ F_{\phi\theta} &= g \sin \theta, \quad \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} r^2 \sin \theta \frac{1}{r^2} \frac{1}{r^2 \sin^2 \theta} g \sin \theta = 0. \end{aligned} \quad (8)$$

Correspondingly, the Dirac equation in this electromagnetic field takes the form

$$\left[i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \gamma^3 \partial_r + \frac{1}{r} \Sigma_{\theta,\phi}^k - M \right] \psi(x) = 0, \quad (9)$$

where

$$\Sigma_{\theta,\phi}^k = i \gamma^1 \partial_\theta + \gamma^2 \frac{i \partial_\phi + (i \sigma^{12} - k) \cos \theta}{\sin \theta}, \quad (10)$$

and $k \equiv eg/hc$. As readily verified, the wave operator in (9) commutes with the following three ones

$$\begin{aligned} J_1^k &= l_1 + \frac{(i \sigma^{12} - k) \cos \phi}{\sin \theta}, \\ J_2^k &= l_2 + \frac{(i \sigma^{12} - k) \sin \phi}{\sin \theta}, \quad J_3^k = l_3, \end{aligned} \quad (11)$$

which obey the $su(2)$ Lie algebra. Clearly, this monopole situation come entirely under the Schrödinger [36] and Pauli [37] approach (detailed treatment of the method was given in [40]). Correspondingly to diagonalizing the \vec{J}_k^2 and J_3^k , the function ψ is to be taken as ($D_\sigma \equiv D_{-m,\sigma}^j(\phi, \theta, 0)$ stands for Wigner functions [38])

$$\psi_{\epsilon jm}^k(t, r, \theta, \phi) = e^{-i\epsilon t} \begin{vmatrix} f_1 & D_{k-1/2} \\ f_2 & D_{k+1/2} \\ f_3 & D_{k-1/2} \\ f_4 & D_{k+1/2} \end{vmatrix}. \quad (12)$$

Further, with the help of recursive relations [38]

$$\begin{aligned} \partial_\theta D_{k+1/2} &= a D_{k-1/2} - b D_{k+3/2}, \quad \partial_\theta D_{k-1/2} = c D_{k-3/2} - a D_{k+1/2}, \\ \sin^{-1} \theta [-m - (k + 1/2) \cos \theta] D_{k+1/2} &= (-a D_{k-1/2} - b D_{k+3/2}), \\ \sin^{-1} \theta [-m - (k - 1/2) \cos \theta] D_{k-1/2} &= (-c D_{k-3/2} - a D_{k+1/2}), \end{aligned}$$

$$\begin{aligned} b &= \frac{\sqrt{(j - k - 1/2)(j + k + 3/2)}}{2}, \\ c &= \frac{\sqrt{(j + k - 1/2)(j - k + 3/2)}}{2}, \\ a &= \frac{1}{2} \sqrt{(j + 1/2)^2 - k^2} \end{aligned}$$

we find how the $\Sigma_{\theta,\phi}^k$ acts on $\psi_{\epsilon jm}^k$

$$\Sigma_{\theta,\phi}^k \psi_{\epsilon jm}^k = i\sqrt{(j+1/2)^2 - k^2} e^{-i\epsilon t} \begin{pmatrix} -f_4 D_{k-1/2} \\ +f_3 D_{k+1/2} \\ +f_2 D_{k-1/2} \\ -f_1 D_{k+1/2} \end{pmatrix}; \quad (13)$$

hereafter the factor $\sqrt{(j+1/2)^2 - k^2}$ will be denoted by ν . For the $f_i(r)$, the radial system derived is

$$\begin{aligned} \frac{\epsilon}{\sqrt{\Phi}} f_3 - i\sqrt{\Phi} \frac{d}{dr} f_3 - i\frac{\nu}{r} f_4 - M f_1 &= 0, \\ \frac{\epsilon}{\sqrt{\Phi}} f_4 + i\sqrt{\Phi} \frac{d}{dr} f_4 + i\frac{\nu}{r} f_3 - M f_2 &= 0, \\ \frac{\epsilon}{\sqrt{\Phi}} f_1 + i\sqrt{\Phi} \frac{d}{dr} f_1 + i\frac{\nu}{r} f_2 - M f_3 &= 0, \\ \frac{\epsilon}{\sqrt{\Phi}} f_2 - i\sqrt{\Phi} \frac{d}{dr} f_2 - i\frac{\nu}{r} f_1 - M f_4 &= 0. \end{aligned} \quad (14)$$

Else one operator can be diagonalized together with $i\partial_t, \bar{J}_k^2, J_3^k$: namely, a generalized Dirac operator

$$\hat{K}^k = -i\gamma^0 \gamma^3 \Sigma_{\theta,\phi}^k. \quad (15)$$

From the equation $\hat{K}^k \psi_{\epsilon jm} = \lambda \psi_{\epsilon jm}$ we find two possible eigenvalues and restrictions on $f_i(r)$

$$f_4 = \delta f_1, \quad f_3 = \delta f_2, \quad \lambda = -\delta \sqrt{(j+1/2)^2 - k^2}. \quad (16)$$

Correspondingly, the system (14) reduces to

$$\begin{aligned} \left(\sqrt{\Phi} \frac{d}{dr} + \frac{\nu}{r} \right) f + \left(\frac{\epsilon}{\sqrt{\Phi}} + \delta M \right) g &= 0, \\ \left(\sqrt{\Phi} \frac{d}{dr} - \frac{\nu}{r} \right) g - \left(\frac{\epsilon}{\sqrt{\Phi}} - \delta M \right) f &= 0, \end{aligned} \quad (17)$$

where

$$f = \frac{f_1 + f_2}{\sqrt{2}}, \quad g = \frac{f_1 - f_2}{\sqrt{2}i}.$$

It is known that quantization of $k = eg/hc$ and j is given by

$$\begin{aligned} eg/hc &= \pm 1/2, \pm 1, \pm 3/2, \dots; \\ j &= |k| - 1/2, |k| + 1/2, |k| + 3/2, \dots \end{aligned} \quad (18)$$

The case of minimal value $j_{min} = |k| - 1/2$ must be treated separately in a special way. For example, let $k = +1/2$, then to the minimal value $j = 0$ there corresponds the wave function in terms of only (t, r) -dependent quantities

$$\psi_{k=+1/2}^{(j=0)}(x) = e^{-i\epsilon t} \begin{pmatrix} f_1(r) \\ 0 \\ f_3(r) \\ 0 \end{pmatrix}. \quad (19)$$

At $k = -1/2$, we have

$$\psi_{k=-1/2}^{(j=0)}(x) = e^{-i\epsilon t} \begin{vmatrix} 0 \\ f_2(r) \\ 0 \\ f_4(r) \end{vmatrix}. \quad (20)$$

Thus, if $k = \pm 1/2$, then to the minimal values j_{\min} there correspond the function substitutions which do not depend at all on the angular variables (θ, ϕ) ; at this point there exists some formal analogy between these electron-monopole states and S -states (with $l = 0$) for a boson field of spin zero: $\Phi_{l=0} = \Phi(r, t)$. However, it would be unwise to attach too much significance to this formal similarity because that (θ, ϕ) -independence of $(e - g)$ -states is not a fact invariant under tetrad gauge transformations.

In contrast, the relation below (let $k = +1/2$)

$$\Sigma_{\theta, \phi}^{+1/2} \psi_{k=+1/2}^{(j=0)}(x) = \gamma^2 \cot \theta (i\sigma^{12} - 1/2) \psi_{k=+1/2}^{(j=0)} \equiv 0 \quad (21)$$

is invariant under arbitrary tetrad gauge transformations. Correspondingly, the matter equation (9) takes on the form

$$\left(i \frac{\gamma^0}{\sqrt{\Phi}} \frac{\partial}{\partial t} + i \gamma^3 \sqrt{\Phi} \frac{\partial}{\partial r} - M \right) \psi^{(j=0)} = 0. \quad (22)$$

It is readily verified that both (19) and (20) representations are directly extended to $(e - g)$ -states with $j = j_{\min}$ at all the other $k = \pm 1, \pm 3/2, \dots$. Indeed,

$$k = +1, +3/2, +2, \dots, \quad \psi_{j_{\min.}}^{k>0}(x) = e^{-i\epsilon t} \begin{vmatrix} f_1(r) D_{k-1/2} \\ 0 \\ f_3(r) D_{k-1/2} \\ 0 \end{vmatrix}; \quad (23)$$

$$k = -1, -3/2, -2, \dots, \quad \psi_{j_{\min.}}^{k<0}(x) = e^{-i\epsilon t} \begin{vmatrix} 0 \\ f_2(r) D_{k+1/2} \\ 0 \\ f_4(r) D_{k+1/2} \end{vmatrix}, \quad (24)$$

and the relation $\Sigma_{\theta, \phi} \psi_{j_{\min}} = 0$ still holds. For instance, let us consider in more detail the case of positive k . Using the recursive relations

$$\begin{aligned} \partial_{\theta} D_{k-1/2} &= \frac{1}{2} \sqrt{2k-1} D_{k-3/2}, \\ \sin^{-1} \theta [-m - (k - 1/2) \cos \theta] D_{k-1/2} &= -\frac{1}{2} \sqrt{2k-1} D_{k-3/2}, \end{aligned}$$

we get

$$\begin{aligned} i\gamma^1 \partial_{\theta} \begin{vmatrix} f_1(r) D_{k-1/2} \\ 0 \\ f_3(r) D_{k-1/2} \\ 0 \end{vmatrix} &= \frac{i}{2} \sqrt{2k-1} \begin{vmatrix} 0 \\ -f_3(r) D_{k-3/2} \\ 0 \\ +f_1(r) D_{k-3/2} \end{vmatrix}, \\ \gamma^2 \frac{i\partial_{\phi} + (i\sigma^{12} - k) \cos \theta}{\sin \theta} \begin{vmatrix} f_1(r) D_{k-1/2} \\ 0 \\ f_3(r) D_{k-1/2} \\ 0 \end{vmatrix} &= \frac{i}{2} \sqrt{2k-1} \begin{vmatrix} 0 \\ +f_3(r) D_{k-3/2} \\ 0 \\ -f_1(r) D_{k-3/2} \end{vmatrix}; \end{aligned}$$

in a sequence, the identity $\Sigma_{\theta,\phi} \psi_{j_{\min}} \equiv 0$ holds. The case of negative k can be considered in the same way. Thus, at every k , the j_{\min} -state equation has the same unique form

$$\left(i \frac{\gamma^0}{\sqrt{\Phi}} \frac{\partial}{\partial t} + i \gamma^3 \sqrt{\Phi} \frac{\partial}{\partial r} - M \right) \psi_{j_{mi}} = 0 ; \quad (25)$$

which leads to the same unique radial system

$$k = +1/2, +1, \dots$$

$$\begin{aligned} \frac{\epsilon}{\sqrt{\Phi}} f_3 - i \sqrt{\Phi} \frac{d}{dr} f_3 - M f_1 &= 0 , \\ \frac{\epsilon}{\sqrt{\Phi}} f_1 + i \sqrt{\Phi} \frac{d}{dr} f_1 - M f_3 &= 0 ; \end{aligned} \quad (26)$$

$$k = -1/2, -1, \dots$$

$$\begin{aligned} \frac{\epsilon}{\sqrt{\Phi}} f_4 + i \sqrt{\Phi} \frac{d}{dr} f_4 - M f_2 &= 0 , \\ \frac{\epsilon}{\sqrt{\Phi}} f_2 - i \sqrt{\Phi} \frac{d}{dr} f_2 - M f_4 &= 0 . \end{aligned} \quad (27)$$

In the limit of flat space-time, these equations are equivalent respectively to $k = +1/2, +1, \dots$

$$\left(\frac{d^2}{dr^2} + \epsilon^2 - m^2 \right) f_1 = 0 , \quad f_3 = \frac{1}{m} \left(\epsilon + i \frac{d}{dr} \right) f_1 ; \quad (28)$$

$$k = -1/2, -1, \dots$$

$$\left(\frac{d^2}{dr^2} + \epsilon^2 - m^2 \right) f_4 = 0 , \quad f_2 = \frac{1}{m} \left(\epsilon + i \frac{d}{dr} \right) f_4 . \quad (29)$$

These equation both lead us to the functions $f = \exp(\pm \sqrt{m^2 - \epsilon^2} r)$. In particular, at $\epsilon < m$, we have a solution

$$\exp (- \sqrt{m^2 - \epsilon^2} r) , \quad (30)$$

which seems to be appropriate to describe bound states in the electron-monopole system.

4 Solution of the radial equations

Let us turn back to the system (17) and (for definiteness) consider equations at $\delta = +1$ (formally the second case $\delta = -1$ corresponds to the change $M \Rightarrow -M$)

$$\begin{aligned} \left(\sqrt{\Phi} \frac{d}{dr} + \frac{\nu}{r} \right) f + \left(\frac{\epsilon}{\sqrt{\Phi}} + M \right) g &= 0 , \\ \left(\sqrt{\Phi} \frac{d}{dr} - \frac{\nu}{r} \right) g - \left(\frac{\epsilon}{\sqrt{\Phi}} - M \right) f &= 0 . \end{aligned} \quad (31)$$

Here we see additional singularities at the points

$$\epsilon + \sqrt{\Phi} M = 0 \quad \text{or} \quad \epsilon - \sqrt{\Phi} M = 0 .$$

For instance, the equation for $f(r)$ has the form

$$\begin{aligned} \frac{d^2}{dr^2} f + \left(\frac{2r}{1+r^2} - \frac{Mr}{\sqrt{1+r^2}(\epsilon + M\sqrt{1+r^2})} \right) \frac{d}{dr} f + \left(\frac{\epsilon^2}{(1+r^2)^2} - \frac{M^2}{1+r^2} \right. \\ \left. - \frac{\nu^2}{r^2(1+r^2)} - \frac{\nu}{r^2(1+r^2)^{3/2}} - \frac{M\nu}{(1+r^2)(\epsilon + M\sqrt{1+r^2})} \right) f = 0 . \end{aligned}$$

However, there exists possibility to move these singularities away through a special transformation of the functions $f(r), g(r)$ [24]. To this end, let us introduce a new variable $r = \sinh \rho$, eqs. (31) look simpler

$$\begin{aligned} \left(\frac{d}{d\rho} + \frac{\nu}{\sinh \rho} \right) f + \left(\frac{\epsilon}{\cosh \rho} + M \right) g &= 0 , \\ \left(\frac{d}{d\rho} - \frac{\nu}{\sinh \rho} \right) g - \left(\frac{\epsilon}{\cosh \rho} - M \right) f &= 0 . \end{aligned} \tag{32}$$

Summing and subtracting two last equations, we get

$$\begin{aligned} \frac{d}{d\rho} (f + g) + \frac{\nu}{\sinh \rho} (f - g) - \frac{\epsilon}{\cosh \rho} (f - g) + M(f + g) &= 0 , \\ \frac{d}{d\rho} (f - g) + \frac{\nu}{\sinh \rho} (f + g) + \frac{\epsilon}{\cosh \rho} (f + g) - M(f - g) &= 0 . \end{aligned} \tag{33}$$

Introducing two new functions

$$f + g = e^{-\rho/2} (F + G) , \quad f - g = e^{+\rho/2} (F - G) , \tag{34}$$

or in matrix form

$$\begin{vmatrix} G \\ H \end{vmatrix} = \begin{vmatrix} \cosh \rho/2 & -\sinh \rho/2 \\ -\sinh \rho/2 & \cosh \rho/2 \end{vmatrix} \begin{vmatrix} g \\ h \end{vmatrix} , \tag{35}$$

where (see definition of the variable z below)

$$\cosh \frac{\rho}{2} = \sqrt{\frac{\sqrt{1-z}+1}{2}} , \quad \sinh \frac{\rho}{2} = \sqrt{\frac{\sqrt{1-z}-1}{2}} , \tag{36}$$

one transforms (33) into

$$\begin{aligned} \frac{d}{d\rho} e^{-\rho/2} (F + G) + \frac{\nu}{\sinh \rho} e^{+\rho/2} (F - G) \\ - \frac{\epsilon}{\cosh \rho} e^{+\rho/2} (F - G) + M e^{-\rho/2} (F + G) &= 0 , \\ \frac{d}{d\rho} e^{+\rho/2} (F - G) + \frac{\nu}{\sinh \rho} e^{-\rho/2} (F + G) \\ + \frac{\epsilon}{\cosh \rho} e^{-\rho/2} (F + G) - M e^{+\rho/2} (F - G) &= 0 , \end{aligned}$$

or

$$\begin{aligned} \frac{d}{d\rho}(F+G) - \frac{1}{2}(F+G) + \frac{\nu}{\sinh \rho}(\cosh \rho + \sinh \rho)(F-G) \\ - \frac{\epsilon}{\cosh \rho}(\cosh \rho + \sinh \rho)(F-G) + M(F+G) = 0, \end{aligned}$$

$$\begin{aligned} \frac{d}{d\rho}(F-G) + \frac{1}{2}(F-G) + \frac{\nu}{\sinh \rho}(\cosh \rho - \sinh \rho)(F+G) \\ + \frac{\epsilon}{\cosh \rho}(\cosh \rho - \sinh \rho)(F+G) - M(F-G) = 0. \end{aligned}$$

Now summing and subtracting two last equations, we obtain

$$\begin{aligned} \left(\frac{d}{d\rho} + \nu \frac{\cosh \rho}{\sinh \rho} - \epsilon \frac{\sinh \rho}{\cosh \rho} \right) F + \left(\epsilon + M - \nu - \frac{1}{2} \right) G = 0, \\ \left(\frac{d}{d\rho} - \nu \frac{\cosh \rho}{\sinh \rho} + \epsilon \frac{\sinh \rho}{\cosh \rho} \right) G + \left(-\epsilon + M + \nu - \frac{1}{2} \right) F = 0. \end{aligned} \quad (37)$$

Let us translate eqs. (37) to the variable z :

$$\begin{aligned} r^2 = \sinh^2 \rho = -z, \quad \frac{d}{d\rho} = 2\sqrt{-z(1-z)} \frac{d}{dz}, \\ \left(2\sqrt{-z(1-z)} \frac{d}{dz} + \nu \frac{\sqrt{1-z}}{\sqrt{-z}} - \epsilon \frac{\sqrt{-z}}{\sqrt{1-z}} \right) F \\ + \left(\epsilon + M - \nu - \frac{1}{2} \right) G = 0, \\ \left(2\sqrt{-z(1-z)} \frac{d}{dz} - \nu \frac{\sqrt{1-z}}{\sqrt{-z}} + \epsilon \frac{\sqrt{-z}}{\sqrt{1-z}} \right) G \\ + \left(-\epsilon + M + \nu - \frac{1}{2} \right) F = 0. \end{aligned} \quad (38)$$

From (38) it follow two 2-nd order differential equations for F and G respectively

$$\begin{aligned} z(1-z) \frac{d^2 F}{dz^2} + \left(\frac{1}{2} - z \right) \frac{dF}{dz} \\ + \left[\frac{1}{4} \left(M - \frac{1}{2} \right)^2 - \frac{\epsilon(\epsilon-1)}{4(1-z)} - \frac{\nu(\nu+1)}{4z} \right] F = 0, \\ z(1-z) \frac{d^2 G}{dz^2} + \left(\frac{1}{2} - z \right) \frac{dG}{dz} \\ + \left[\frac{1}{4} \left(M - \frac{1}{2} \right)^2 - \frac{\epsilon(\epsilon+1)}{4(1-z)} - \frac{\nu(\nu-1)}{4z} \right] G = 0. \end{aligned} \quad (39)$$

With the use of substitutions

$$F = z^A (1-z)^B \bar{F}(z), \quad G = z^K (1-z)^L \bar{G}(z),$$

eqs. (39) take the form

$$\begin{aligned}
& z(1-z) \frac{d^2 \bar{F}}{dz^2} + \left[2A + \frac{1}{2} - (2A + 2B + 1)z \right] \frac{d\bar{F}}{dz} \\
& + \left[\frac{1}{4} \left(M - \frac{1}{2} \right)^2 - (A + B)^2 - \frac{\epsilon(\epsilon - 1) + 2B(1 - 2B)}{4(1 - z)} \right. \\
& \quad \left. - \frac{\nu(\nu + 1) - 2A(2A - 1)}{4z} \right] \bar{F} = 0,
\end{aligned} \tag{40}$$

$$\begin{aligned}
& z(1-z) \frac{d^2 \bar{G}}{dz^2} + \left[2K + \frac{1}{2} - (2K + 2L + 1)z \right] \frac{d\bar{G}}{dz} \\
& + \left[\frac{1}{4} \left(M - \frac{1}{2} \right)^2 - (K + L)^2 - \frac{\epsilon(\epsilon + 1) + 2L(1 - 2L)}{4(1 - z)} \right. \\
& \quad \left. - \frac{\nu(\nu - 1) - 2K(2K - 1)}{4z} \right] \bar{G} = 0.
\end{aligned} \tag{41}$$

First let us consider eq. (40); at A and B taken accordingly

$$A = \frac{1 + \nu}{2}, \quad -\frac{\nu}{2}, \quad B = \frac{\epsilon}{2}, \quad \frac{1 - \epsilon}{2} \tag{42}$$

it becomes simpler

$$\begin{aligned}
& z(1-z) \frac{d^2 f}{dz^2} + \left[2A + \frac{1}{2} - (2A + 2B + 1)z \right] \frac{df}{dz} \\
& + \left[\frac{1}{4} \left(M - \frac{1}{2} \right)^2 - (A + B)^2 \right] f = 0,
\end{aligned} \tag{43}$$

which is of hypergeometric type with parameters

$$a = \frac{M}{2} - \frac{1}{4} + A + B, \quad b = -\frac{M}{2} + \frac{1}{4} + A + B, \quad c = 2A + 1/2.$$

To construct functions appropriate to describe bound states we must choose

$$A = \frac{1 + \nu}{2} > 0, \quad B = \frac{1 - \epsilon}{2} < 0, \quad c = \nu + 3/2; \tag{44}$$

polynomial solutions will arise with the quantization rule imposed

$$\begin{aligned}
a &= -n, \quad \epsilon_n = M + 2n + \nu + \frac{3}{2}, \\
b &= -n - M - 1/2, \quad c = \nu + 3/2.
\end{aligned} \tag{45}$$

Now let us turn to eq. (41). At A, B chosen according to

$$K = \frac{1 - \nu}{2}, \quad \frac{\nu}{2}, \quad L = -\frac{\epsilon}{2}, \quad \frac{1 + \epsilon}{2} \tag{46}$$

it will be simpler

$$z(1-z) \frac{d^2 g}{dz^2} + \left[2K + \frac{1}{2} - (2K + 2L + 1)z \right] \frac{dg}{dz} + \left[\frac{1}{4} \left(M - \frac{1}{2} \right)^2 - (K + L)^2 \right] g = 0, \quad (47)$$

which is of hypergeometric type

$$\alpha = \frac{M}{2} - \frac{1}{4} + K + L, \quad \beta = -\frac{M}{2} + \frac{1}{4} + K + L, \quad \gamma = 2K + \frac{1}{2}.$$

Again, to get bound states we choose the values

$$K = \frac{\nu}{2} > 0, \quad L = -\frac{\epsilon}{2} < 0, \quad (48)$$

then the quantization rule arises

$$\alpha = -N, \quad \epsilon_N = M + 2N + \nu - \frac{1}{2}. \quad (49)$$

It can be noted that $\epsilon_N = \epsilon_n$, when $N = n + 1$.

Let us calculate relative coefficient between functions $F(z)$ and $G(z)$. These being taken in the form

$$F(z) = F_0 z^{(1+\nu)/2} (1-z)^{(1-\epsilon)/2} \bar{F}(a, b, c; z), \quad c = \frac{3}{2} + \nu, \\ a = \frac{M}{2} + \frac{3}{4} + \frac{\nu}{2} - \frac{\epsilon}{2}, \quad b = -\frac{M}{2} + \frac{5}{4} + \frac{\nu}{2} - \frac{\epsilon}{2}; \quad (50)$$

and

$$G(z) = G_0 z^{\nu/2} (1-z)^{-\epsilon/2} \bar{G}(\alpha, \beta, \gamma; z), \quad \gamma = \frac{1}{2} + \nu = c - 1, \\ \alpha = \frac{M}{2} - \frac{1}{4} + \frac{\nu}{2} - \frac{\epsilon}{2} = a - 1, \quad \beta = -\frac{M}{2} + \frac{1}{4} + \frac{\nu}{2} - \frac{\epsilon}{2} = b - 1, \quad (51)$$

must obey the following system

$$\left(2\sqrt{-z(1-z)} \frac{d}{dz} + \nu \frac{\sqrt{1-z}}{\sqrt{-z}} - \epsilon \frac{\sqrt{-z}}{\sqrt{1-z}} \right) F + (+\epsilon + M - \nu - \frac{1}{2}) G = 0, \\ \left(2\sqrt{-z(1-z)} \frac{d}{dz} - \nu \frac{\sqrt{1-z}}{\sqrt{-z}} + \epsilon \frac{\sqrt{-z}}{\sqrt{1-z}} \right) G + (-\epsilon + M + \nu - \frac{1}{2}) F = 0.$$

To find a relative factor, it is convenient to use the second equation

$$\left(-2\sqrt{-z(1-z)} \frac{d}{dz} - \nu \frac{\sqrt{1-z}}{\sqrt{-z}} + \epsilon \frac{\sqrt{-z}}{\sqrt{1-z}} \right) G + (-\epsilon + M + \nu - \frac{1}{2}) F = 0.$$

Substituting expressions for F and G , after simple calculation we get to

$$2iG_0 \frac{d\bar{G}}{dz} = F_0 \left(-\epsilon + M + \nu - \frac{1}{2}\right) \bar{F}.$$

Allowing for the known rule for differentiating hypergeometric functions

$$\begin{aligned} \frac{d}{dz} \bar{G}(z) &= \frac{d}{dz} F(a-1, b-1, c-1; z) \\ &= \frac{(a-1)(b-1)}{c-1} F(a, b, c; z) = \frac{(a-1)(b-1)}{c-1} \bar{F}(z), \end{aligned}$$

we obtain

$$2iG_0 \frac{(a-1)(b-1)}{c-1} = F_0 \left(-\epsilon + M + \nu - \frac{1}{2}\right).$$

Ultimately, we arrive at the formula

$$F_0 = i \frac{M - 1/2 + N}{2} G_0, \quad (52)$$

remembering that $\epsilon_N = M - 1/2 + 2N + \nu$.

5 Radial equations in the case j_{min}

Let us turn back to the case of the minimal value of j :

$$k = +1/2, +1, \dots$$

$$\begin{aligned} \frac{\epsilon}{\sqrt{\Phi}} f_3 - i \sqrt{\Phi} \frac{d}{dr} f_3 - M f_1 &= 0, \\ \frac{\epsilon}{\sqrt{\Phi}} f_1 + i \sqrt{\Phi} \frac{d}{dr} f_1 - M f_3 &= 0; \end{aligned} \quad (53)$$

from where for new functions

$$H = \frac{f_1 + f_3}{\sqrt{2}}, \quad G = \frac{f_1 - f_3}{i\sqrt{2}}$$

we derive

$$k = +1/2, +1, \dots$$

$$\begin{aligned} \sqrt{\Phi} \frac{d}{dr} H + \left(\frac{\epsilon}{\sqrt{\Phi}} + M\right) G &= 0, \\ \sqrt{\Phi} \frac{d}{dr} G - \left(\frac{\epsilon}{\sqrt{\Phi}} - M\right) H &= 0. \end{aligned} \quad (54)$$

And in the same manner for another case we have

$$k = -1/2, -1, \dots$$

$$\begin{aligned} \frac{\epsilon}{\sqrt{\Phi}} f_4 + i \sqrt{\Phi} \frac{d}{dr} f_4 - M f_2 &= 0, \\ \frac{\epsilon}{\sqrt{\Phi}} f_2 - i \sqrt{\Phi} \frac{d}{dr} f_2 - M f_4 &= 0, \end{aligned} \quad (55)$$

from whence for new functions

$$H = \frac{f_2 + f_4}{\sqrt{2}}, \quad G = \frac{f_2 - f_4}{i\sqrt{2}}$$

we obtain

$$\begin{aligned} \sqrt{\Phi} \frac{d}{dr} G + \left(\frac{\epsilon}{\sqrt{\Phi}} - M \right) H &= 0, \\ \sqrt{\Phi} \frac{d}{dr} H - \left(\frac{\epsilon}{\sqrt{\Phi}} + M \right) G &= 0. \end{aligned} \quad (56)$$

We can use the above method to eliminate nonphysical singular points. Let us perform special transformation on the functions

$$G + H = e^{-\rho/2}(g + h), \quad G - H = e^{+\rho/2}(g - h). \quad (57)$$

After simple calculation we arrive at

instead of (54)

$$\begin{aligned} \left(\frac{d}{d\rho} + \epsilon \frac{\sinh \rho}{\cosh \rho} \right) g + (-\epsilon + M - 1/2) h &= 0, \\ \left(\frac{d}{d\rho} - \epsilon \frac{\sinh \rho}{\cosh \rho} \right) h + (+\epsilon + M - 1/2) g &= 0; \end{aligned} \quad (58)$$

instead of (56)

$$\begin{aligned} \left(\frac{d}{d\rho} + \epsilon \frac{\sinh \rho}{\cosh \rho} \right) h + (-\epsilon - M - 1/2) g &= 0, \\ \left(\frac{d}{d\rho} - \epsilon \frac{\sinh \rho}{\cosh \rho} \right) g + (+\epsilon - M - 1/2) h &= 0. \end{aligned} \quad (59)$$

In the variable z

$$r = \sinh \rho = \sqrt{-z}$$

the system (58) takes the form

$$\begin{aligned} \sqrt{-z(1-z)} \left(\frac{d}{dz} - \frac{\epsilon/2}{1-z} \right) g - \frac{(-\epsilon + M - 1/2)}{2} h &= 0, \\ \sqrt{-z(1-z)} \left(\frac{d}{dz} + \frac{\epsilon/2}{1-z} \right) h - \frac{(+\epsilon + M - 1/2)}{2} g &= 0. \end{aligned} \quad (60)$$

Note that the system is symmetric with respect to changes

$$f \Longleftrightarrow h, \quad \epsilon \Longleftrightarrow -\epsilon. \quad (61)$$

After excluding the function h from (60) we get

$$h = \frac{2}{(-\epsilon + M - 1/2)} \sqrt{(-z)(1-z)} \left(\frac{d}{dz} - \frac{\epsilon/2}{1-z} \right) g ,$$

$$\sqrt{(-z)(1-z)} \left(\frac{d}{dz} + \frac{\epsilon/2}{1-z} \right) \sqrt{(-z)(1-z)} \left(\frac{d}{dz} - \frac{\epsilon/2}{1-z} \right) g$$

$$- \frac{(M - 1/2)^2 - \epsilon^2}{4} g = 0 . \quad (62)$$

Ultimately, an equation for $g(z)$ reads

$$z(1-z) \frac{d^2 g}{dz^2} + (1/2 - z) \frac{dg}{dz} + \left(\frac{(M - 1/2)^2}{4} - \frac{\epsilon^2 + \epsilon}{4} \frac{1}{1-z} \right) g = 0 . \quad (63)$$

In the same manner we get a second order differential equation for h after exclusion of g :

$$g = \frac{2}{(+\epsilon + M - 1/2)} \sqrt{(-z)(1-z)} \left(\frac{d}{dz} + \frac{\epsilon/2}{1-z} \right) h ,$$

$$\sqrt{(-z)(1-z)} \left(\frac{d}{dz} - \frac{\epsilon/2}{1-z} \right) \sqrt{(-z)(1-z)} \left(\frac{d}{dz} + \frac{\epsilon/2}{1-z} \right) h$$

$$- \frac{(M - 1/2)^2 - \epsilon^2}{4} h = 0 , \quad (64)$$

and ultimately

$$z(1-z) \frac{d^2 h}{dz^2} + (1/2 - z) \frac{dh}{dz} + \left(\frac{(M - 1/2)^2}{4} - \frac{\epsilon^2 - \epsilon}{4} \frac{1}{1-z} \right) h = 0 . \quad (65)$$

Equations (65) and (63) differ only in the sign at the parameter ϵ .

6 Solutions of radial equations in the case j_{min}

With the use of substitution $g = (1-z)^A \varphi(z)$, from (63) we produce for φ

$$z(1-z) \varphi'' + \left[\frac{1}{2} - (1+2A)z \right] \varphi'$$

$$+ \left[\left(A^2 - \frac{A}{2} - \frac{\epsilon^2 + \epsilon}{4} \right) \frac{1}{1-z} - A^2 + \frac{(M - 1/2)^2}{4} \right] \varphi . \quad (66)$$

Requiring

$$A^2 - \frac{A}{2} - \frac{\epsilon^2 + \epsilon}{4} = 0 \quad \implies \quad 2A = \epsilon + 1, -\epsilon$$

one gets

$$z(1-z) \varphi'' + \left[\frac{1}{2} - (1+2A)z \right] \varphi' - \frac{4A^2 - (M - 1/2)^2}{4} \varphi = 0 ,$$

$$\varphi = F(a, b, c, z) , \quad c = \frac{1}{2} , \quad a + b = 2A, \quad ab = \frac{4A^2 - (M - 1/2)^2}{4} , \quad (67)$$

that is

$$a = \frac{2A + (M - 1/2)}{2}, \quad b = \frac{2A - (M - 1/2)}{2}. \quad (68)$$

Below we will use negative values for A

$$A = -\epsilon/2, \quad g(z) = (1 - z)^{-\epsilon/2} \varphi(z); \quad (69)$$

so that

$$a = \frac{-\epsilon + (M - 1/2)}{2}, \quad b = \frac{-\epsilon - (M - 1/2)}{2}. \quad (70)$$

Any 2-nd order differential equation has two linearly independent solutions; here they are

$$\varphi_1 = U_1(z) = F(a, b, c; z),$$

$$\varphi_2 = U_5(z) = z^{1-c} F(a + 1 - c, b + 1 - c, 2 - c; z). \quad (71)$$

Similar analysis can be performed for eq. (65)

$$z(1 - z) \frac{d^2 h}{dz^2} + (1/2 - z) \frac{dh}{dz} + \left(\frac{(M - 1/2)^2}{4} - \frac{\epsilon^2 - \epsilon}{4} \frac{1}{1 - z} \right) h = 0. \quad (72)$$

With the use of substitution $h(z) = (1 - z)^L \eta(z)$, for $\eta(z)$ we produce

$$z(1 - z) \eta'' + \left[\frac{1}{2} - (1 + 2L)z \right] \eta' + \left[\left(L^2 - \frac{L}{2} - \frac{\epsilon^2 - \epsilon}{4} \right) \frac{1}{1 - z} - L^2 + \frac{(M - 1/2)^2}{4} \right] \eta = 0. \quad (73)$$

Requiring

$$L^2 - \frac{L}{2} - \frac{\epsilon^2 - \epsilon}{4} = 0 \quad \implies \quad 2L = +\epsilon, -\epsilon + 1$$

one gets

$$\begin{aligned} z(1 - z) \eta'' + \left[\frac{1}{2} - (1 + 2L)z \right] \eta' - \frac{4L^2 - (M - 1/2)^2}{4} \eta &= 0, \\ \eta &= F(\alpha, \beta, \gamma, z), \quad \gamma = \frac{1}{2}, \\ \alpha + \beta &= 2L, \quad \alpha\beta = \frac{4L^2 - (M - 1/2)^2}{4}, \end{aligned} \quad (74)$$

that is

$$\alpha = \frac{2L - (M - 1/2)}{2}, \quad \beta = \frac{2L + (M - 1/2)}{2}. \quad (75)$$

Below we will use negative values for L

$$L = (-\epsilon + 1)/2 < 0, \quad h(z) = (1 - z)^{(-\epsilon+1)/2} \eta(z), \quad (76)$$

so that

$$\alpha = \frac{-\epsilon + 1 + (M - 1/2)}{2}, \quad \beta = \frac{-\epsilon + 1 - (M - 1/2)}{2}. \quad (77)$$

Functions $g(z)$ and $h(z)$ must obey the above system of first order differential equations. To verify that, let us start with the functions

$$\begin{aligned} g &= G_0(1 - z)^A \varphi_1(z), & 2A &= -\epsilon, \\ \varphi_1 &= F(a, b, c, z) & c &= 1/2, \\ a &= \frac{-\epsilon + (M - 1/2)}{2}, & b &= \frac{-\epsilon - (M - 1/2)}{2}, \end{aligned} \quad (78)$$

$$\begin{aligned} h &= H_0(1 - z)^L \eta_2(z), & 2L &= -\epsilon + 1, \\ \eta_2 &= z^{1/2} F(\alpha + 1 - \gamma, \beta + 1 - \gamma, 2 - \gamma, z), \\ \alpha + 1 - \gamma &= \frac{-\epsilon + 2 + (M - 1/2)}{2} = a + 1, \\ \beta + 1 - \gamma &= \frac{-\epsilon + 2 - (M - 1/2)}{2} = b + 1, \\ 2 - \gamma &= c + 1, \end{aligned} \quad (79)$$

and relate them with the help of the first equation in the system (60)

$$\sqrt{(-z)(1 - z)} \left(\frac{d}{dz} - \frac{\epsilon/2}{1 - z} \right) g - \frac{(-\epsilon + M - 1/2)}{2} h = 0.$$

After simple calculations we obtain

$$G_0 i \frac{d}{dz} F(a, b, c, z) = H_0 \frac{(-\epsilon + M - 1/2)}{2} F(a + 1, b + 1, c + 1, z),$$

from whence it follows

$$G_0 i \frac{ab}{c} = H_0 \frac{(-\epsilon + M - 1/2)}{2},$$

that is

$$H_0 = i(-\epsilon - M + 1/2) G_0.$$

To get polynomial solutions we must require

$$\begin{aligned} a &= -n \implies \epsilon_n = M + 2n - 1/2, \\ b &= -n - M + 1/2, & c &= 1/2, \\ g(z) &= (1 - z)^{-(\epsilon_n+1)/2} F(a, b, c, z). \end{aligned} \quad (80)$$

note that

$$g(z) = (1 - z)^{-n-(M-1/2)/2} {}_1F(-n, -n - M + \frac{1}{2}, \frac{1}{2}, z); \quad (81)$$

therefore as $z = -r^2 \longrightarrow -\infty$ the function $g(z)$ tends to zero

$$g(z) \longrightarrow 0 , \quad \text{only if} \quad M > \frac{1}{2} .$$

In usual units, that condition for existence of bound states consistent with anti de Sitter geometry structure, the inequality $M > \frac{1}{2}$, looks as

$$\rho > \frac{1}{2} \frac{\hbar}{Mc} = \frac{1}{2} \lambda_e = 1.213 \times 10^{-12} \text{ metre}$$

so it can be broken only in a very strong anti de Sitter gravitation background, the latter is beyond of our treatment.

Let us write down several energy levels (in usual units)

$$\epsilon_0 = Mc^2 - \frac{1}{2} \frac{c\hbar}{\rho} , \quad \epsilon_1 = Mc^2 + \frac{3}{2} \frac{c\hbar}{\rho} , \quad \epsilon_2 = Mc^2 + \frac{5}{2} \frac{c\hbar}{\rho} , \dots \quad (82)$$

or

$$\epsilon_0 = Mc^2 \left(1 - \frac{1}{2} \frac{\lambda_e}{\rho}\right) , \quad \epsilon_1 = Mc^2 \left(1 + \frac{3}{2} \frac{\lambda_e}{\rho}\right) , \quad \epsilon_2 = Mc^2 \left(1 + \frac{5}{2} \frac{\lambda_e}{\rho}\right) , \dots \quad (83)$$

If one mentally increases the curvature radius ρ , the energy levels will become denser and the minimal level tends to the value Mc^2

$$\epsilon_0 = Mc^2 \left(1 - \frac{1}{2} \frac{\lambda_e}{\rho}\right) \longrightarrow Mc^2 . \quad (84)$$

7 Conclusions and discussion

To understand better results, let us discuss the case of minimal j_{min} in the limit of vanishing curvature. To this end, let us specify in more detail solutions for minimal values j_{min} in Minkowski space:

$$k = +1/2, +1, \dots$$

$$\begin{aligned} \epsilon f_3 - i \frac{d}{dr} f_3 - M f_1 &= 0 , \\ \epsilon f_1 + i \frac{d}{dr} f_1 - M f_3 &= 0 ; \end{aligned} \quad (85)$$

$$k = -1/2, -1, \dots$$

$$\begin{aligned} \epsilon f_4 + i \frac{d}{dr} f_4 - M f_2 &= 0 , \\ \epsilon f_2 - i \frac{d}{dr} f_2 - M f_4 &= 0 . \end{aligned} \quad (86)$$

Let detail the case of positive $k = +1/2, +1, \dots$. Let it be

$$\frac{f_1 + f_3}{\sqrt{2}} = h(r) , \quad \frac{f_1 - f_3}{i\sqrt{2}} = g(r) \quad (87)$$

relevant equations are

$$\frac{d}{dr}h + (\epsilon + M)g = 0 , \quad \frac{d}{dr}g - (\epsilon - M)h = 0 . \quad (88)$$

With the substitutions

$$h(r) = He^{\gamma r} , \quad g(r) = Ge^{\gamma r} \quad (89)$$

we get (first let it be $(\epsilon^2 - M^2) > 0$)

$$\begin{aligned} \gamma^2 = -(\epsilon^2 - M^2) &\equiv -p^2 , & \gamma = +ip, -ip . \\ H\gamma + (\epsilon + M)G = 0 &\quad \text{or} \quad G\gamma - (\epsilon - M)H = 0 . \end{aligned} \quad (90)$$

Thus we have two linearly independent solutions

$$\begin{aligned} h_1(r) &= H_1 e^{+ipr} , & g_1(r) &= G_1 e^{+ipr} , & G_1 &= \frac{\epsilon - M}{ip} H_1 ; \\ h_2(r) &= H_2 e^{-ipr} , & g_2(r) &= G_2 e^{-ipr} , & G_2 &= \frac{\epsilon - M}{-ip} H_2 . \end{aligned} \quad (91)$$

Below, we take $H_1 = H_2 = 1$. We can introduce two linear combinations of these solutions the first

$$\begin{aligned} \frac{h_1(r) + h_2(r)}{2} &= \cos pr , \\ \frac{g_1(r) + g_2(r)}{2} &= \frac{\epsilon - M}{p} \sin pr ; \end{aligned} \quad (92)$$

the second

$$\begin{aligned} \frac{h_1(r) - h_2(r)}{2i} &= \sin pr , \\ \frac{g_1(r) - g_2(r)}{2i} &= \frac{\epsilon - M}{-p} \cos pr . \end{aligned} \quad (93)$$

Now let us specify the case $(\epsilon^2 - M^2) < 0$

$$\begin{aligned} \gamma^2 = -(\epsilon^2 - M^2) &\equiv +q^2 , & \gamma &= +q, -q . \\ H\gamma + (\epsilon + M)G = 0 &\quad \text{or} \quad G\gamma - (\epsilon - M)H = 0 . \end{aligned} \quad (94)$$

Thus we have two linearly independent solutions

$$\begin{aligned} h_1(r) &= H_1 e^{+qr} , & g_1(r) &= G_1 e^{+qr} , & G_1 &= \frac{\epsilon - M}{q} H_1 ; \\ h_2(r) &= H_2 e^{-qr} , & g_2(r) &= G_2 e^{-qr} , & G_2 &= \frac{\epsilon - M}{-q} H_2 . \end{aligned} \quad (95)$$

Below, we take $H_1 = H_2 = 1$. We can introduce two linear combinations of these solutions the first

$$\begin{aligned}\frac{h_1(r) + h_2(r)}{2} &= \cosh qr , \\ \frac{g_1(r) + g_2(r)}{2} &= \frac{\epsilon - M}{q} \sinh qr\end{aligned}\tag{96}$$

the second

$$\begin{aligned}\frac{h_1(r) - h_2(r)}{2} &= \sinh qr , \\ \frac{g_1(r) - g_2(r)}{2} &= \frac{\epsilon - M}{q} \cosh qr .\end{aligned}\tag{97}$$

Evidently, above constructed solutions in de Sitter model provide us with generalizations of these of Minkowski space. It may be verified additionally by direct limiting process when $\rho \rightarrow \infty$. To this end, let us translate solutions in de Sitter space to usual units

$$\begin{aligned}g_1(R) &= \left(1 + \frac{R^2}{\rho^2}\right)^{-\frac{E\rho}{2c\hbar}} F(a, b, c; -\frac{R^2}{\rho^2}) , \quad c = 1/2 , \\ g_2(R) &= R \left(1 + \frac{R^2}{\rho^2}\right)^{-\frac{E\rho}{2c\hbar}} F(a + 1 - c, b + 1 - c, 2 - c; -\frac{R^2}{\rho^2}) , \\ h_1(R) &= \left(1 + \frac{R^2}{\rho^2}\right)^{-\frac{E\rho}{2c\hbar} + 1/2} F(\alpha, \beta, \gamma; -\frac{R^2}{\rho^2}) , \quad \gamma = 1/2 , \\ h_2(R) &= R \left(1 + \frac{R^2}{\rho^2}\right)^{-\frac{E\rho}{2c\hbar} + 1/2} F(\alpha + 1 - \gamma, \beta + 1 - \gamma, 2 - \gamma; -\frac{R^2}{\rho^2}) ,\end{aligned}$$

Parameters of hypergeometric functions are given by

$$\begin{aligned}a &= \frac{1}{2} \left(-\frac{E\rho}{c\hbar} + \left(\frac{mc\rho}{\hbar} - \frac{1}{2}\right) \right) , \quad b = \frac{1}{2} \left(-\frac{E\rho}{c\hbar} - \left(\frac{mc\rho}{\hbar} - \frac{1}{2}\right) \right) , \\ \alpha &= \frac{1}{2} \left(-\frac{E\rho}{c\hbar} + 1 + \left(\frac{mc\rho}{\hbar} - \frac{1}{2}\right) \right) , \quad \beta = \frac{1}{2} \left(-\frac{E\rho}{c\hbar} + 1 - \left(\frac{mc\rho}{\hbar} - \frac{1}{2}\right) \right) .\end{aligned}$$

Let us examine the limiting procedure at $\rho \rightarrow \infty$ in $F(a, b, c; -R^2/\rho^2)$. Because

$$\begin{aligned}\frac{1}{1!} \frac{ab}{c} \left(-\frac{R^2}{\rho^2}\right) &\rightarrow \frac{1}{2!} (m^2 c^2 / \hbar^2 - E^2 / \hbar^2 c^2) R^2 = -\frac{1}{2!} (pR)^2 , \\ \frac{1}{2!} \frac{a(a+1)b(b+1)}{c(c+1)} \left(-\frac{R^2}{\rho^2}\right) &\rightarrow +\frac{(pR)^4}{4!} , \\ \frac{1}{3!} \frac{a(a+1)(a+2)b(b+1)(b+2)}{c(c+1)(c+2)} \left(-\frac{R^2}{\rho^2}\right) &\rightarrow -\frac{(pR)^6}{6!} ,\end{aligned}$$

and so on, we obtain the following limiting relation

$$\lim_{\rho \rightarrow \infty} F(a, b, c; -\frac{R^2}{\rho^2}) = \cos pr \quad \implies \quad \lim_{\rho \rightarrow \infty} g_1(R) = \cos pr .$$

Similarly, we get

$$\lim_{\rho \rightarrow \infty} h_1(R) = \cos pR . \quad (98)$$

In the same manner, we arrive at two limiting relationships

$$\lim_{\rho \rightarrow \infty} pR g_2(R) = \sin pR , \quad \lim_{\rho \rightarrow \infty} pR h_2(R) = \sin pR . \quad (99)$$

To rationalize how the finite sums (polynomials of n -order) may approximate the functions $\cos pR$ and $\sin pR$ (infinite series), we should take into account the quantization condition

$$\alpha = -n \quad \Longrightarrow \quad E = Mc^2 + (2n - \frac{1}{2}) \frac{c\hbar}{\rho}$$

At any fixed E , as ρ increases the number n also must increase. This means, that the finite sums of n terms when ρ increases will approximate infinite series.

8 Acknowledgements

Authors are grateful to V.M. Red'kov for moral support and advices. This work was supported by the Fund for Basic Researches of Belarus F11M-152.

References

- [1] Gibbons, G.W.: Anti-de-Sitter spacetime and its uses, in Mathematical and quantum aspects of relativity and cosmology (Pythagoreon, 1998), Lecture Notes in Phys., 537, Springer-Verlag, 102-142, 2000.
- [2] Dirac, P.A.M.: The electron wave equation in the de Sitter space. Ann. Math. **36**, 657–669 (1935)
- [3] Dirac, P.A.M.: Wave equations in conformal space. Ann. of Math. **37**, 429–442 (1936)
- [4] Schrödinger, E.: The proper vibrations of the expanding universe. Physica. **6**, 899–912 (1939)
- [5] Schrödinger, E.: General theory of relativity and wave mechanics. Wiss. en Natuurkund. **10**, 2–9 (1940)
- [6] Goto, K.: Wave equations in de Sitter space. Progr. Theor. Phys. **6**, 1013–1014 (1951)
- [7] Nachtmann, O.: Quantum theory in de-Sitter space. Commun. Math. Phys. **6**, 1–16 (1967)
- [8] Chernikov, N.A., Tagirov, E.A.: Quantum theory of scalar field in de Sitter space-time. Ann. Inst. Henri Poincare. **IX**, 109–141 (1968)
- [9] Börner, G., Dürr, H.P.: Classical and quantum theory in de Sitter space. Nuovo Cim. A. **64**, 669–713 (1969)

- [10] Fushchych, W.L., Krivsky, I.Yu.: On representations of the inhomogeneous de Sitter group and equations in five-dimensional Minkowski space. Nucl. Phys. B. **14**, 573–585 (1969)
- [11] Börner, G., Dürr, H.P.: Classical and Quantum Fields in de Sitter space. Nuovo Cim. **LXIV**, 669 (1969)
- [12] Castagnino, M.: Champs spinoriels en Relativité générale; le cas particulier de l’espace-temps de De Sitter et les équations d’ond pour les spins élevés. Ann. Inst. Henri Poincaré. A. **16**, 293–341 (1972)
- [13] Tagirov, E.A.: Consequences of field quantization in de Sitter type cosmological models. Ann. Phys. **76**, 561–579 (1973)
- [14] S.W. Hawking, G.F.R. Ellis. The large scale structure of space-time. Cambridge University Press, 1973.
- [15] Riordan, F.: Solutions of the Dirac equation in finite de Sitter space. Nuovo Cim. B. **20**, 309–325 (1974)
- [16] Candelas, P., Raine, D.J.: General-relativistic quantum field theory: an exactly soluble model. Phys. Rev. D. **12**, 965–974 (1975)
- [17] Schomblond, Ch., Spindel P.: Propagateurs des champs spinoriels et vectoriels dans l’univers de de Sitter. Bull. Cl. Sci., V. Ser., Acad. R. Belg. **LXII**, 124 (1976)
- [18] Hawking, S.W., Gibbons, G.W.: Cosmological event horizons, thermodynamics, and particle creation. Phys. Rev. D. **15**, 2738–2751 (1977)
- [19] Avis, S.J., Isham, C.J., Storey, D.: Quantum Field Theory In Anti-de Sitter Space-Time. Phys. Rev. D. **18**, 3565 (1978)
- [20] S.J. Avis, C.J. Isham, D. Storey. Quantum field theory in anti-de Sitter space-time. Phys. Rev. D, **18** , 3565–3576 (1978)
- [21] Lohiya, D., Panchapakesan, N.: Massless scalar field in a de Sitter universe and its thermal flux. J. Phys. A. **11**, 1963–1968 (1978)
- [22] Lohiya, D., Panchapakesan, N.: Particle emission in the de Sitter universe for massless fields with spin. J. Phys. A. **12**, 533–539 (1979)
- [23] Hawking, S., Page, D.: Thermodynamics Of Black Holes In Anti-de Sitter Space. Commun. Math. Phys. **87**, 577–588 (1983)
- [24] Otchik, V.S.: On the Hawking radiation of spin 1/2 particles in the de Sitter space-time. Class. Quantum Crav. **2**, 539–543 (1985)
- [25] Motolla, F.: Particle creation in de Sitter space. Phys. Rev. D. **31**, 754–766 (1985)
- [26] Takashi Mishima, Akihiro Nakayama: Particle production in de Sitter spacetime. Progr. Theor. Phys. **77**, 218–222 (1987)

- [27] Polarski, D.: The scalar wave equation on static de Sitter and anti-de Sitter spaces. *Class. Quantum Grav.* **6**, 893–900 (1989)
- [28] Bros, J., Gazeau, J.P, Moschella, U.: Quantum Field Theory in the de Sitter Universe. *Phys. Rev. Lett.* **73**, 1746 (1994)
- [29] Suzuki, H., Takasugi, E.: Absorption Probability of De Sitter Horizon for Massless Fields with Spin. *Mod. Phys. Lett. A.* **11**, 431–436 (1996)
- [30] Pol'shin, S.A.: Group Theoretical Examination of the Relativistic Wave Equations on Curved Spaces. I. Basic Principles. <http://arxiv.org/abs/gr-qc/9803091>; II. De Sitter and Anti-de Sitter Spaces. <http://arxiv.org/abs/gr-qc/9803092>; III. Real reducible spaces. <http://arxiv.org/abs/gr-qc/9809011>
- [31] I.I. Cotaescu. Normalized energy eigenspinors of the Dirac field on anti-de Sitter spacetime. *Phys. Rev. D*(3) **60**, 124006, 4pp (1999)
- [32] Garidi, T., Huguet, E., Renaud, J.: De Sitter Waves and the Zero Curvature Limit Comments. *Phys. Rev. D.* **67**, 124028 (2003)
- [33] Moradi, S., Rouhani, S., Takook, M.V.: Discrete Symmetries for Spinor Field in de Sitter Space. *Phys. Lett. B.* **613**, 74–82 (2005)
- [34] A. Bachelot. The Dirac equation on the Anti-de-Sitter Universe. L'équation de Dirac sur l'univers Anti-de Sitter *Comptes Rendus Mathématique.* **345**, Issue 8, 435-440 (2007)
- [35] V.I. Strazhev, L.M. Tomil'chik. Electrodynamics with a magnetic charge. – Minsk: Nauka i Technika, 1975.
- [36] Schrödinger, E.: The ambiguity of the wave function. *Annalen der Physik.* **32**, 49–55 (1938)
- [37] Pauli, W.: Über die Kriterium für Ein-oder Zweiwertigkeit der Eigenfunktionen in der Wellenmechanik. *Helv. Phys. Acta.* 1939. **12**, 147–168 (1939)
- [38] Varshalovich, D.A., Moskalev, A.N., Hersonskiy, V.K.: Quantum theory of angular momentum. Nauka, Leningrad (1975)
- [39] V.M. Red'kov. Fields in Riemannian space and the Lorentz group. Publishing House "Belarusian Science", Minsk, 496 pages (2009).
- [40] V.M. Red'kov. Tetrad formalism, spherical symmetry and Schrödinger basis. Publishing House "Belarusian Science", Minsk, 339 pages (2011).